

## Fe-Cr Mixed Binary Spinels as Accessory Magnetic Minerals in the Sudetic Ophiolitic Rocks

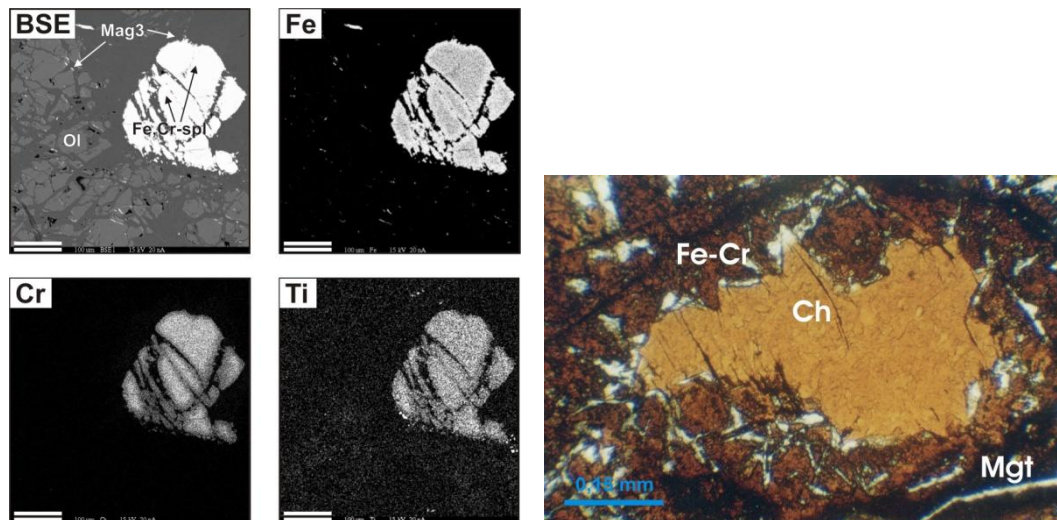
Magdalena KĄDZIAŁKO-HOFMOKL<sup>1</sup> and Tomasz WERNER<sup>1,✉</sup>

<sup>1</sup>Institute of Geophysics, Polish Academy of Sciences, Warsaw, Poland

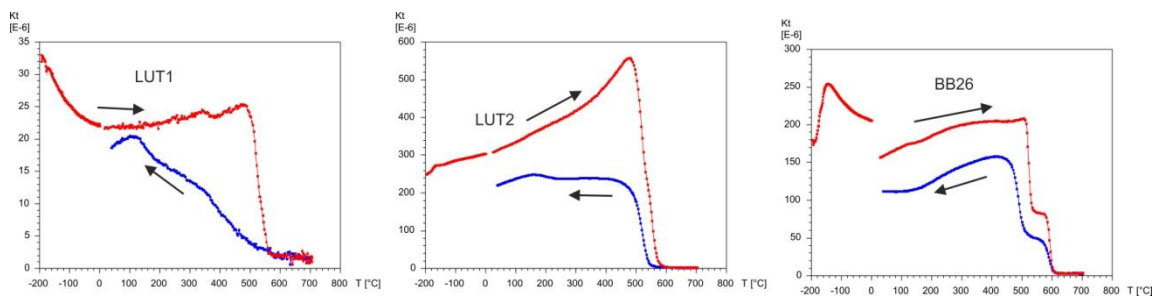
✉ twerner@igf.edu.pl

### Abstract

Sudetic ophiolite is formed of three serpentinite massives situated around the Sowie Góry Mts Block: Jordanów-Gogołów Massif (JGSM), Braszowice-Brzeźnica Massif (BBSM) and Szklary Massif (SZM). Chromite Fe-Cr spinels occur in chromium ore (chromitites) in JGSM and BBSM (ore fragments on waste heaps) and as scattered grains in ultramafic rocks. The Fe-Cr chromite series have a general formula of  $(\text{Fe}^{2+}_{1-x}\text{Fe}^{3+}_x)[\text{Fe}^{2+}_{1-x}\text{Fe}^{3+}_{2-2y-x}\text{Cr}^{3+}_{2y}]_2\text{O}_4$  is built of mixed spinels with end members: primary chromite  $(\text{Fe}^{2+})[\text{Cr}_3^{2+}]\text{O}_4$  ( $y = 0$ ) with normal ordered spinel structure ( $x = 0$ ) and magnetite  $(\text{Fe}^{3+})[\text{Fe}^{2+}\text{Fe}^{3+}\text{O}_4]$  ( $y = 1$ ) with inversed ordered spinel structure ( $x = 1$ ). Fe-Cr spinels are called ferrichromites for  $0.31 < y < 0.63$  and Cr-magnetites for  $y > 0.63$ . The composition affects substantially Curie temperatures:  $T_c$  of primary chromite is  $-202^\circ\text{C}$ ,  $T_c$  above r.t. for ferrichromites up to  $T_c$  of  $585^\circ\text{C}$  for magnetite. The primary chromites crystallize from mafic melt in upper mantle-lower crust environments. They are very stable against metamorphism and retain primary composition of their cores long during later metamorphism. Under cooling below ca  $600^\circ\text{C}$  chromite begin to alter: magnetite starts to replace chromite, with subsolidus exsolutions and oxidations processes. The core of the grain retains its primary composition with typical  $T_c$  of  $-180^\circ$  up  $-120^\circ\text{C}$ , around it ring 1 composed of Fe-Cr solution grains of ferrichromite and ring 2 of Cr-magnetite formed during metamorphism were observed (Fig. 1). Apart of changes in composition additional alterations, namely order – disorder transformation takes place (e.g., described by Harrison and Putnis (1999)). Such transformation is caused by electron hopping between tetrahedral and octahedral sublattices with help of oxygen ions ( $x$  in the above formula presents a fraction of 3+ cations in tetrahedral sublattice). Chromites were studied with magnetic methods as determinations of the magnetic susceptibility upon temperature curves (km-T) at the range  $-190^\circ\text{C}$  up to  $700^\circ\text{C}$  for fresh and previously heated samples as well as with hysteresis properties studies. The Km-T experiments for fresh samples showed a wide spectrum of Km(T) curves depending on the composition (Fig. 2). Changes in composition influence changes in  $T_c$  observed on the heating branch of km-T curve, changes in ordering impart changes in  $T_c$  observed on the cooling branch. Therefore k-Tc curves are irreversible and  $T_c$  of samples observed in Cr-magnetites during heating is higher by  $10\text{--}20^\circ\text{C}$  than observed during cooling. During next heating – cooling cycles both branches become reversible. The observed thermal hysteresis



**Fig. 1.** BSE-SEM maps for chromite grains from chromite ore (JGSM, Tapadła, ferri-chromite).



**Fig. 2.** Magnetic susceptibility Km-T curves (-190°C up to 700°C) for chromites of different composition.

during heating-cooling cycle is due to kinetic lag in cation ordering during cooling (Harrison and Putnis 1996). The coercivity usually increases due to heating suggesting subsolidus exsolutions.

**Keywords:** chromites, magnetic susceptibility, ophiolite.

## References

- Harrison, R., and A. Putnis (1996), Magnetic properties of the magnetite-spinel solid solution: Curie temperatures, magnetic susceptibilities and cation ordering, *Am. Mineral.* **81**, 375–384.
- Harrison, R., and A. Putnis (1999), The magnetic properties and crystal chemistry of oxide spinel solid solutions, *Surv. Geophys.* **19**, 6, 461–520, DOI:10.1023/A:1006535023784.